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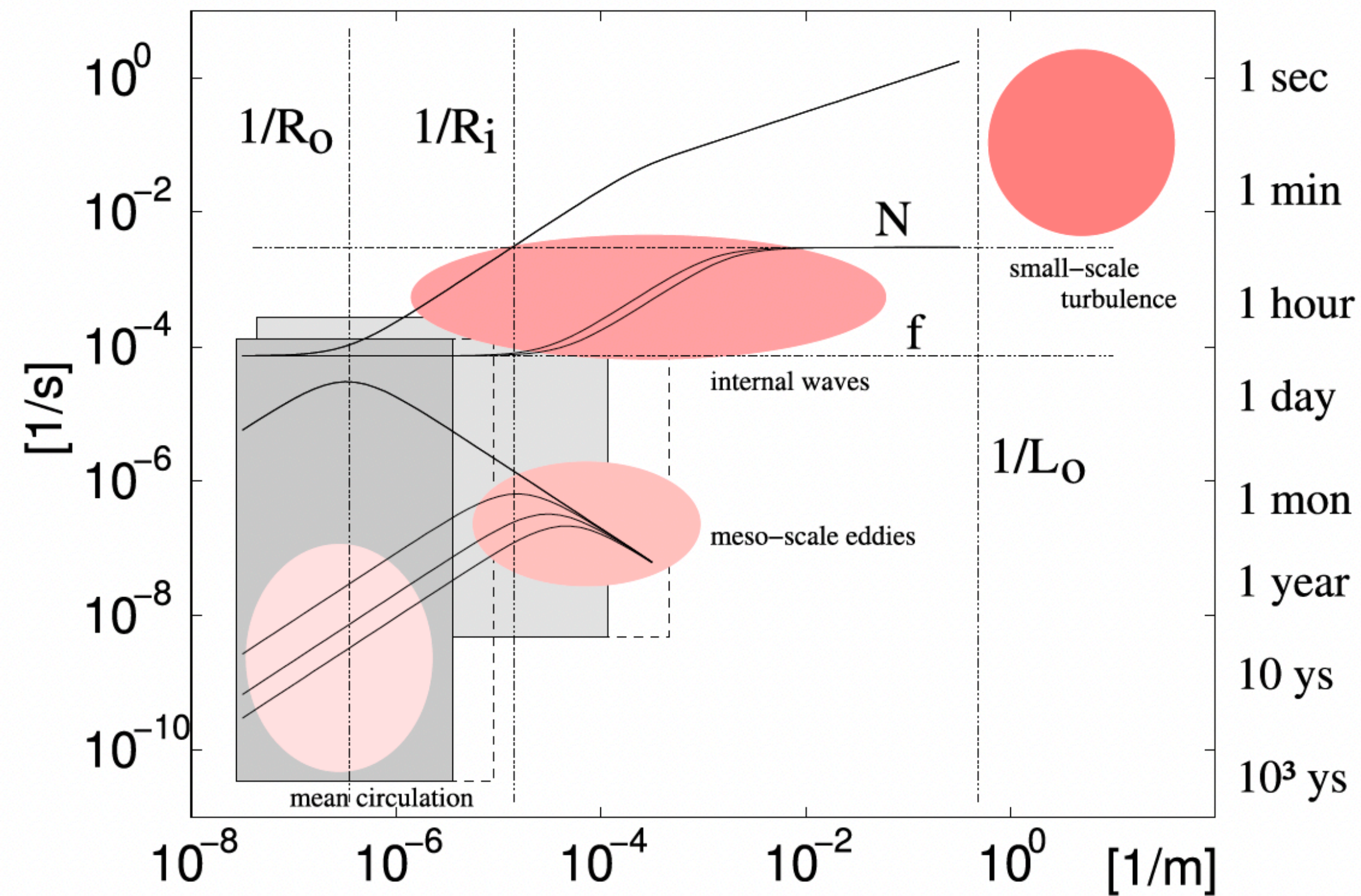
TOWARDS ENERGETICALLY CONSISTENT COUPLING

Reading, 12.04.2024



Collaborative research center TRR181:

- Improve climate predictions by removing spurious energy sources and sinks and provide energetically consistent parameterisations



Eden et al. (2014)

next step: coupling to surface wave model

- identify energy transfers
- provide a simple, but energetically consistent coupling framework



$$\partial_t \mathbf{u} + (\mathbf{u}^L \cdot \nabla) \mathbf{u} + \mathbf{f} \times \mathbf{u}^L = b\mathbf{z} - \nabla p - u_i^L \nabla u_i^S + \mathbf{D}_u$$

$$\partial_t \mathbf{u}^L + (\mathbf{u}^L \cdot \nabla) \mathbf{u}^L + \mathbf{f} \times \mathbf{u}^L = b\mathbf{z} - \nabla p - \mathbf{u}^L \times (\nabla \times \mathbf{u}^S) + \mathbf{D}_u + \partial_t \mathbf{u}^S$$

$\mathbf{u}^L = \mathbf{u} + \mathbf{u}^S$, with Stokes drift \mathbf{u}^S from irrotational waves.

e.g. Craik & Leibovich (1976), Leibovich (1980), Holm (1996), Suzuki & Fox-Kemper (2016)

$$MKE_L = \frac{1}{2}(\bar{\mathbf{u}} + \bar{\mathbf{u}}^S) \cdot (\bar{\mathbf{u}} + \bar{\mathbf{u}}^S) = MKE_E + MKE_S + MKE_{ES}$$

$$MKE_L = \frac{1}{2} \bar{\mathbf{u}}^L \cdot \bar{\mathbf{u}}^L$$

$$MKE_E = \frac{1}{2} \bar{\mathbf{u}} \cdot \bar{\mathbf{u}}$$

$$MKE_S = \frac{1}{2} \bar{\mathbf{u}}^S \cdot \bar{\mathbf{u}}^S$$

$$MKE_{ES} = \bar{\mathbf{u}} \cdot \bar{\mathbf{u}}^S$$



$$\frac{\partial}{\partial t} MKE_E + \bar{u}_j \frac{\partial}{\partial x_j} MKE_L + \frac{\partial}{\partial x_j} (\bar{u}_j \bar{p}) + \frac{\partial}{\partial x_j} (\bar{u}_i \overline{u'_i u'_j}) + \frac{\partial}{\partial x_j} \left(\mu \bar{u}_i \frac{\partial \bar{u}_i^L}{\partial x_j} \right) =$$

$$\overline{u'_i u'_j} \frac{\partial \bar{u}_i}{\partial x_j} + \delta_{i,3} \bar{b} \bar{u}_i - \mu \frac{\partial \bar{u}_i}{\partial x_j} \frac{\partial \bar{u}_i^L}{\partial x_j} + \epsilon_{ijk} \bar{u}_i f_j \bar{u}_k^S$$

$\epsilon_{ijk} \bar{u}_i f_j \bar{u}_k^S$, work done by Coriolis-Stokes (Hasselmann) force is an energy source

,e.g. Liu et al. (2009) , Sayol et al. (2016), Suzuki&Fox-Kemper (2016), Zhang et al. (2019)



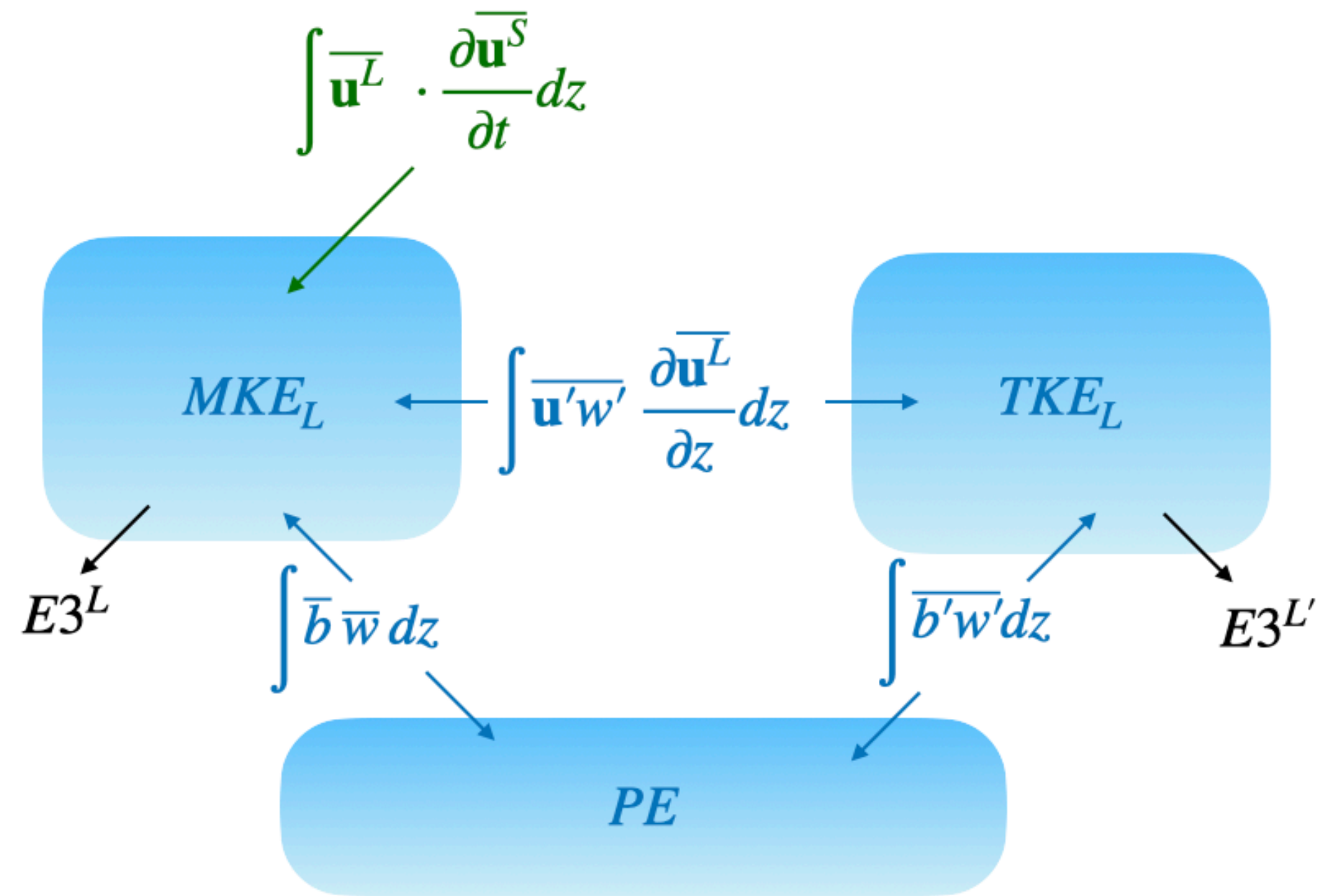
$$\frac{\partial}{\partial t} MKE_L + \overline{u_j^L} \frac{\partial}{\partial x_j} MKE_L + \frac{\partial}{\partial x_j} (\overline{u_j^L} \overline{p}) + \frac{\partial}{\partial x_j} (\overline{u_i^L} \overline{u_i' u_j'}) + \frac{\partial}{\partial x_j} \left(\mu \frac{\partial}{\partial x_j} MKE_L \right) =$$

$$\overline{u_i' u_j'} \frac{\partial \overline{u_i^L}}{\partial x_j} + \delta_{i,3} \overline{b} \overline{u_i^L} - \mu \frac{\partial \overline{u_i^L}}{\partial x_j} \frac{\partial \overline{u_i^L}}{\partial x_j} + \overline{u_i^L} \frac{\partial \overline{u_i^S}}{\partial t}$$

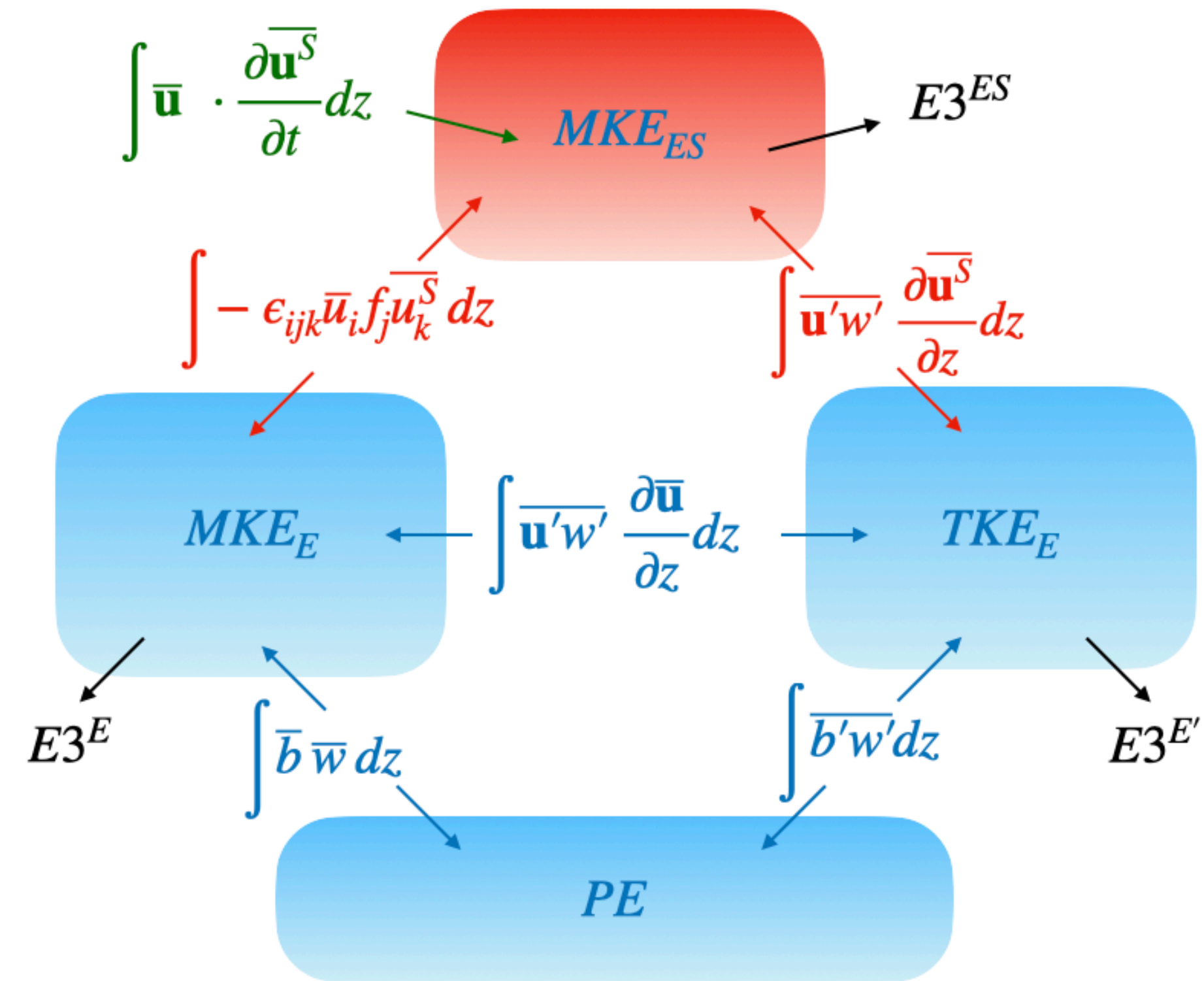
$$\frac{\partial}{\partial t} TKE_L + \overline{u_j^L} \frac{\partial}{\partial x_j} TKE_L + \frac{\partial}{\partial x_j} (\overline{u_j'} \overline{p'}) + \frac{\partial}{\partial x_j} (\overline{u_i' u_i' u_j'}) + \frac{\partial}{\partial x_j} \left(\mu \frac{\partial}{\partial x_j} TKE_L \right) =$$

$$-\overline{u_i' u_j'} \frac{\partial \overline{u_i^L}}{\partial x_j} + \delta_{i,3} \overline{b' u_i'} - \mu \frac{\partial \overline{u_i'}}{\partial x_j} \frac{\partial \overline{u_i'}}{\partial x_j}$$





Lagrangian energy budget



Eulerian energy budget



$$\partial_t \mathbf{u}^L + (\mathbf{u}^L \cdot \nabla) \mathbf{u}^L + \mathbf{f} \times \mathbf{u}^L = b\mathbf{z} - \nabla p - \mathbf{u}^L \times (\nabla \times \mathbf{u}^S) + \mathbf{D}_u + \partial_t \mathbf{u}^S$$

consider a linear and non-viscous ocean away from lateral boundaries, with $\mathbf{u}^S = \begin{pmatrix} u^S(z, t) \\ 0 \\ 0 \end{pmatrix}$

horizontal momentum equations become:

$$\partial_t u^L - f v^L = \partial_t u^S$$

$$\partial_t v^L + f u^L = 0$$

Hasselmann (1970)

and in steady state:

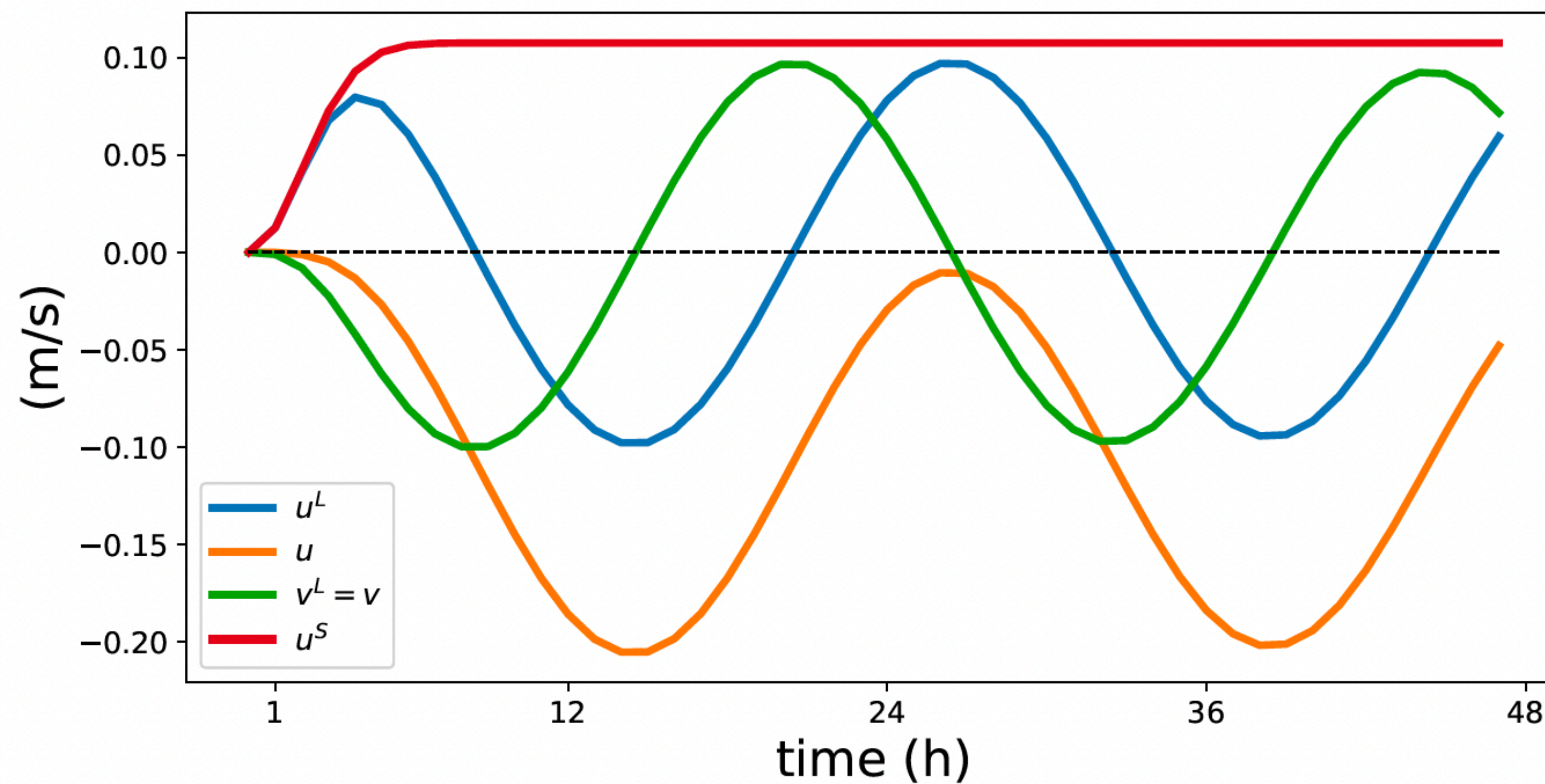
$$u^S = -u$$

$$v = 0$$

Ursell & Deacon (1950), Pollard (1970)



$$u^S(z, t) = u_0^S \exp\left(\frac{z}{D_s}\right) \left[1 - \exp\left(\frac{-t^2}{2T_w^2}\right) \right], \text{ with growth time scale } T_w \text{ of 2h}$$



$$\partial_t u^L - fv = + \partial_t u^S$$

$$\partial_t v^L + fu + fu^S = 0$$

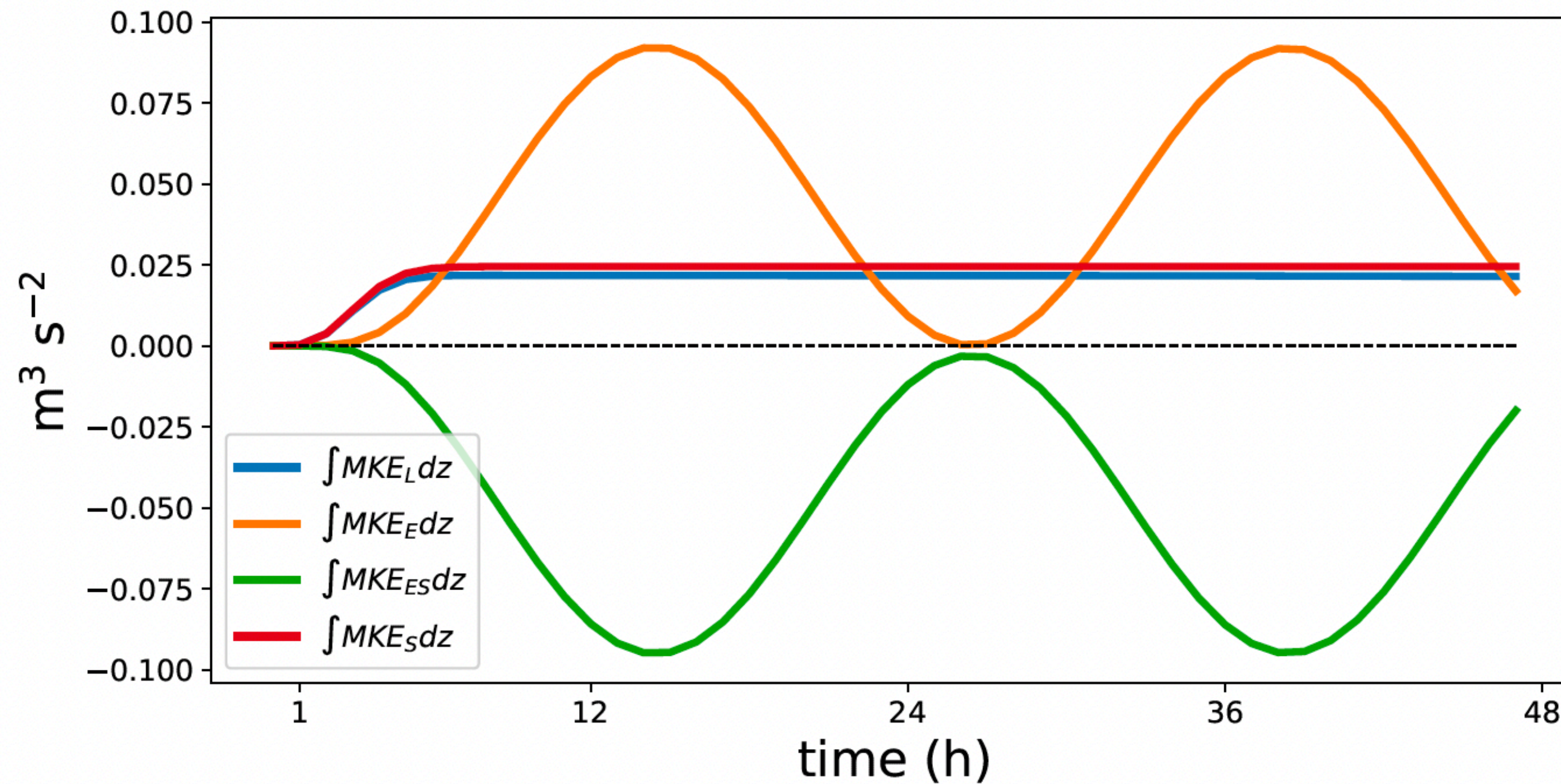


● $\frac{\partial}{\partial t} MKE_L = \bar{\mathbf{u}}^L \cdot \partial_t \bar{\mathbf{u}}^S$

● $\frac{\partial}{\partial t} MKE_E = -\bar{\mathbf{u}} \cdot f\mathbf{z} \times \bar{\mathbf{u}}^S$

● $\frac{\partial}{\partial t} MKE_S = \bar{\mathbf{u}}^S \cdot \partial_t \bar{\mathbf{u}}^S$

● $\frac{\partial}{\partial t} MKE_{ES} = \bar{\mathbf{u}} \cdot f\mathbf{z} \times \bar{\mathbf{u}}^S + \bar{\mathbf{u}} \cdot \partial_t \bar{\mathbf{u}}^S$



$$\partial_t \mathbf{u}^L + (\mathbf{u}^L \cdot \nabla) \mathbf{u}^L + \mathbf{f} \times \mathbf{u}^L = b\mathbf{z} - \nabla p - \mathbf{u}^L \times (\nabla \times \mathbf{u}^S) + \nabla \mu_t \nabla (\mathbf{u}^L - \mathbf{u}^S) + \partial_t \mathbf{u}^S$$

assume $\nabla_h u^L \gg \nabla_h u^S$ and $w^S = 0$

equations to be used in large-scale ocean models become:

$$\partial_t \mathbf{u}_h^L + (\mathbf{u}^L \cdot \nabla) \mathbf{u}_h^L + \mathbf{f} \times \mathbf{u}_h^L = -\nabla_h p + \nabla \mu_t \nabla (\mathbf{u}_h^L - \mathbf{u}_h^S) + \partial_t \mathbf{u}_h^S$$

$$\partial_z p - b = [-u^L \partial_z u^S - v^L \partial_z v^S]$$

$$\partial_t b + (\mathbf{u}^L \cdot \nabla) b = \nabla \kappa_t \nabla b$$

$$\nabla \cdot \mathbf{u}^L = 0$$



$$\partial_t F + \partial_{\mathbf{x}}(\mathbf{v}_g F) = S_{in} + S_{diss} + S_{nl} \quad F(\omega, \theta)$$

$$\Phi_{in} = \rho_w g \int_0^{2\pi} \int_0^{\infty} d\omega d\theta S_{in}$$

$$\Phi_{diss} = \rho_w g \int_0^{2\pi} \int_0^{\infty} d\omega d\theta S_{diss}$$

$$\boldsymbol{\tau}_{in} = \rho_w g \int_0^{2\pi} \int_0^{\infty} d\omega d\theta \frac{\mathbf{k}}{\omega} S_{in}$$

$$\boldsymbol{\tau}_{diss} = \rho_w g \int_0^{2\pi} \int_0^{\infty} d\omega d\theta \frac{\mathbf{k}}{\omega} S_{diss}$$

$$\boldsymbol{\tau}_{oc} = \boldsymbol{\tau}_a - \boldsymbol{\tau}_{in} + \boldsymbol{\tau}_{diss}$$

Chalikov & Belevich (1993), Janssen (2012), Breivik et al. (2015)



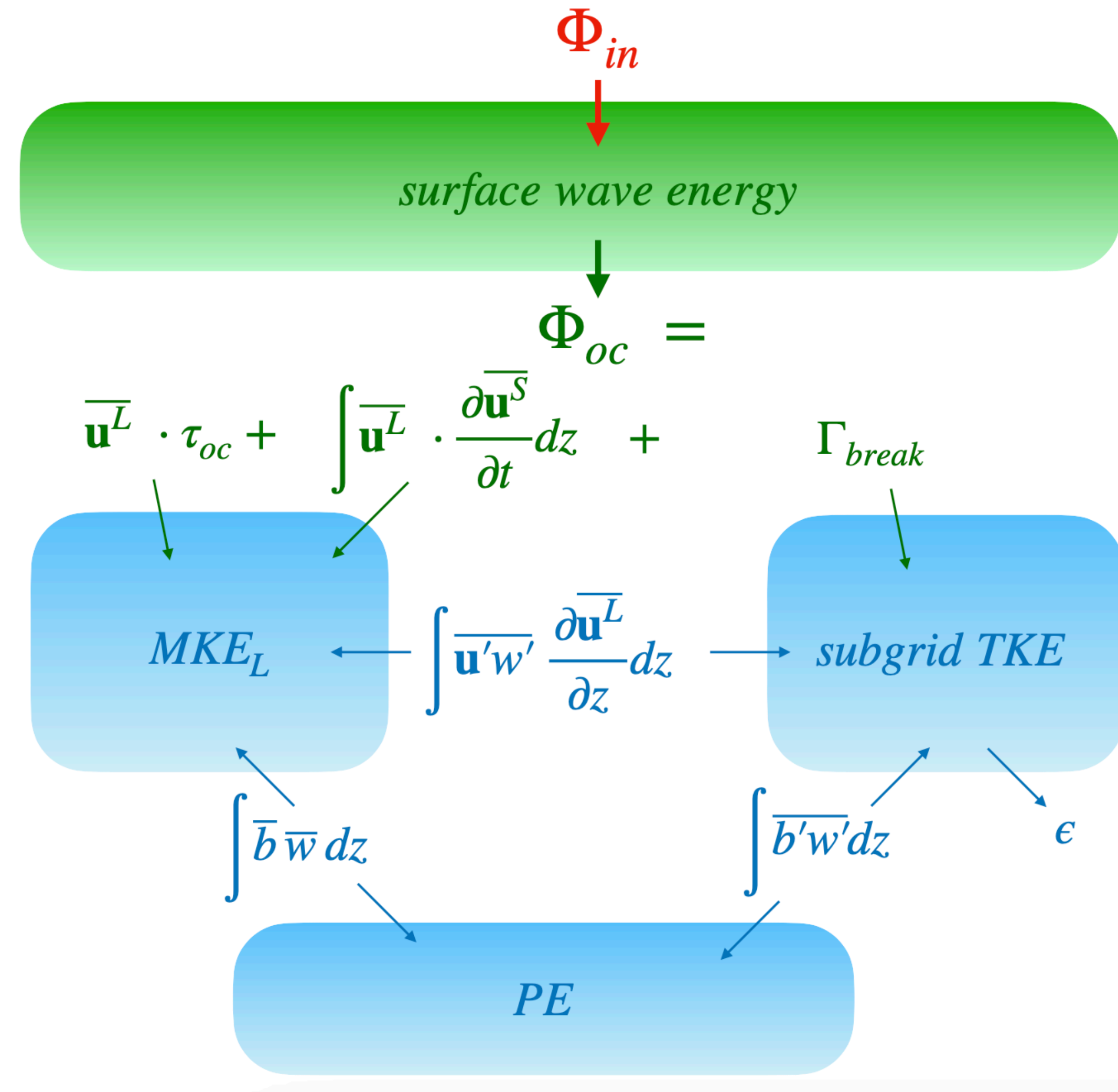
$\omega > \omega_c$ *diagnostic range*

$\omega < \omega_c$ *prognostic range*

$$\Phi_{oc} = \rho_w g \int_0^{2\pi} \int_{\omega_c}^{\infty} S_{in} d\omega d\theta - \Phi_{diss}$$

$$\Phi_{oc} = \overline{\mathbf{u}^L} \cdot \tau_{oc} + \int \overline{\mathbf{u}^L} \cdot \partial_t \overline{\mathbf{u}^S} + \Gamma_{break}$$





- simple inclusion of sea state impacts in climate models by re-interpreting existing velocity as Lagrangian velocity ($\partial_t \mathbf{u}_h^S$ is only new term in momentum equation)

- energy provided by the wave model is split-up into energy which goes into mean motions and a remainder which goes into turbulence according to

$$\Phi_{oc} = \overline{\mathbf{u}^L} \cdot \tau_{oc} + \int \overline{\mathbf{u}^L} \cdot \partial_t \overline{\mathbf{u}^S} + \Gamma_{break}$$

- energy transfer terms in Lagrangian budget are easier to interpret, for example the Coriolis-Stokes term is absent

